



Beyond Einstein: From the Big Bang to Black Holes

LISA Technology Development at NASA/GSFC

J.I. Thorpe

37th COSPAR Scientific Assembly

Montréal, Québec

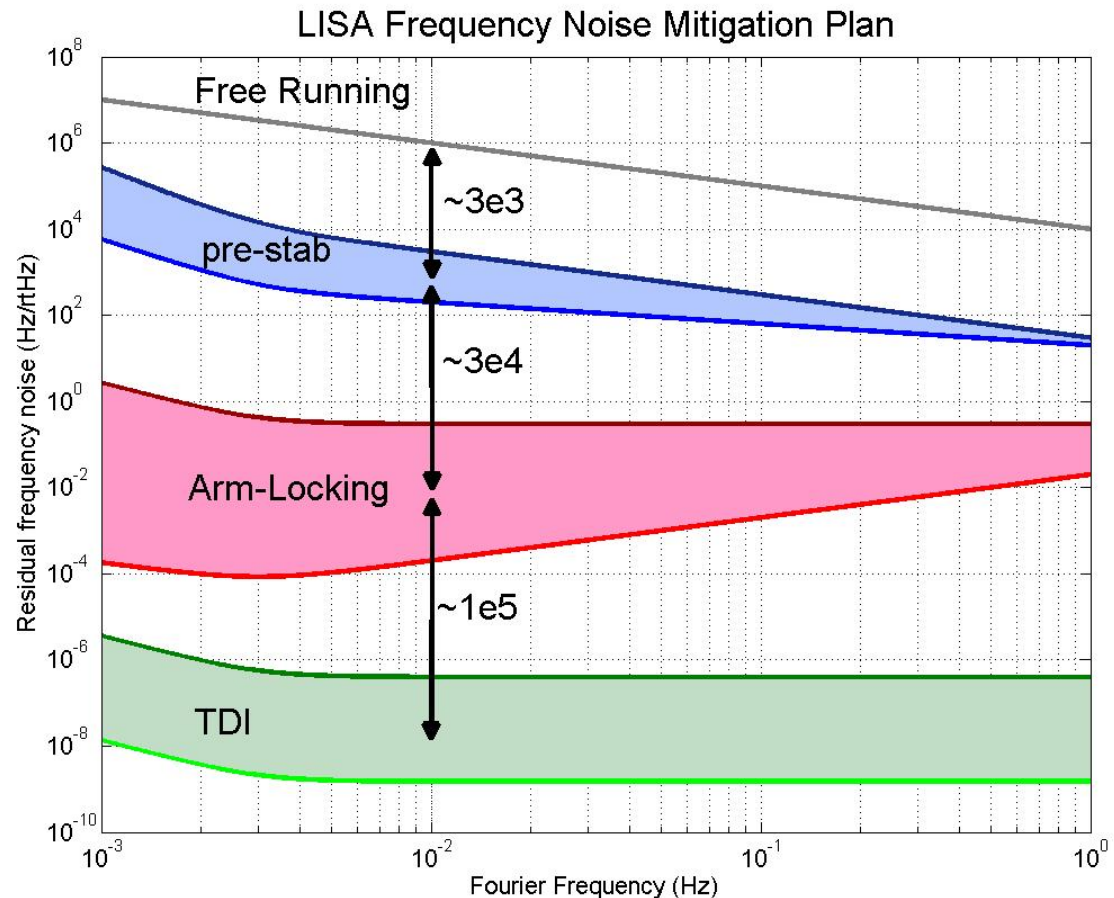
July 16th, 2008

- 🌀 Laser Frequency Stabilization
 - Optical Cavities with frequency tuning
 - Molecular Iodine
- 🌀 Stable Environments
 - Stable test-bed for formation flying
 - Fused-silica fibers for torsion pendula
- 🌀 Surface Effects
 - Kelvin Probe
- 🌀 Laser Study

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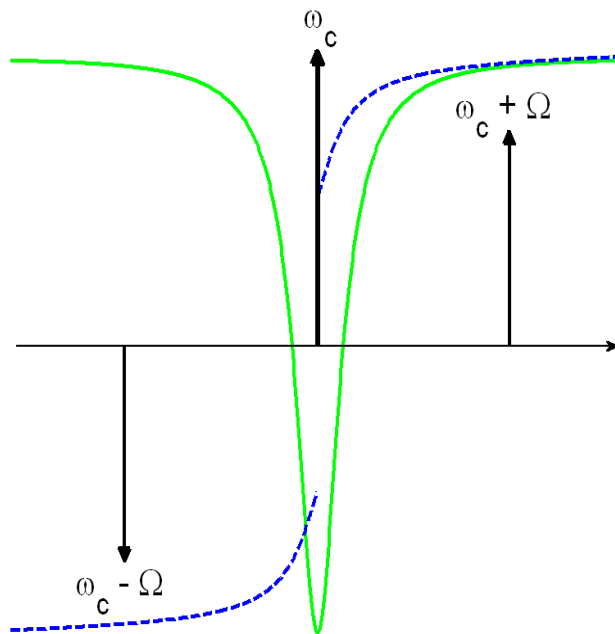
Laser frequency noise is a major potential noise source for LISA

- Three-stage system (two active one passive) to achieve overall suppression of $\sim 10^{13}$
- Running pre-stabilization and arm-locking in series reduces gain (bandwidth) requirements on arm-locking.
- Serial arrangement *requires frequency-tunable pre-stabilization*

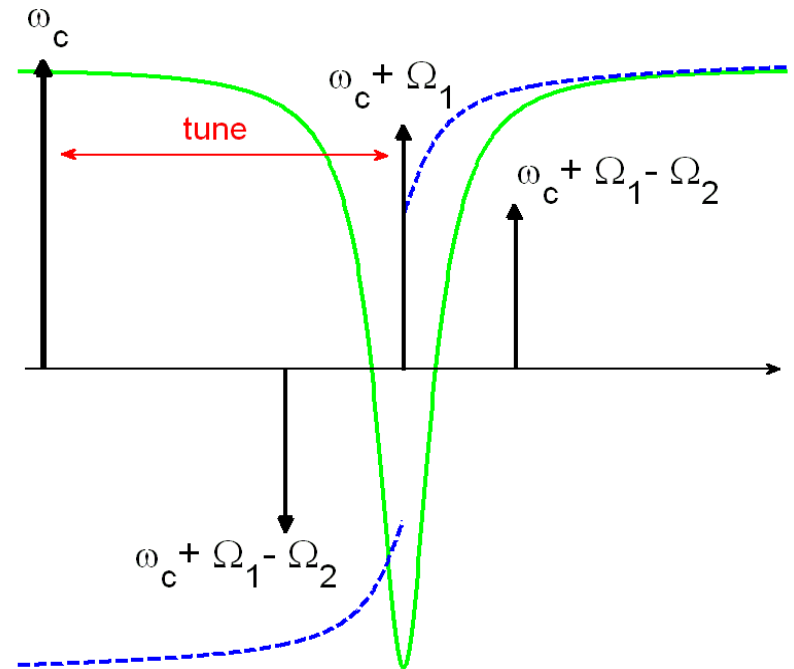


Concept: Lock phase-modulation sidebands to cavity resonance and tune central frequency by adjusting modulation frequency.

Normal Pound-Drever-Hall Lock

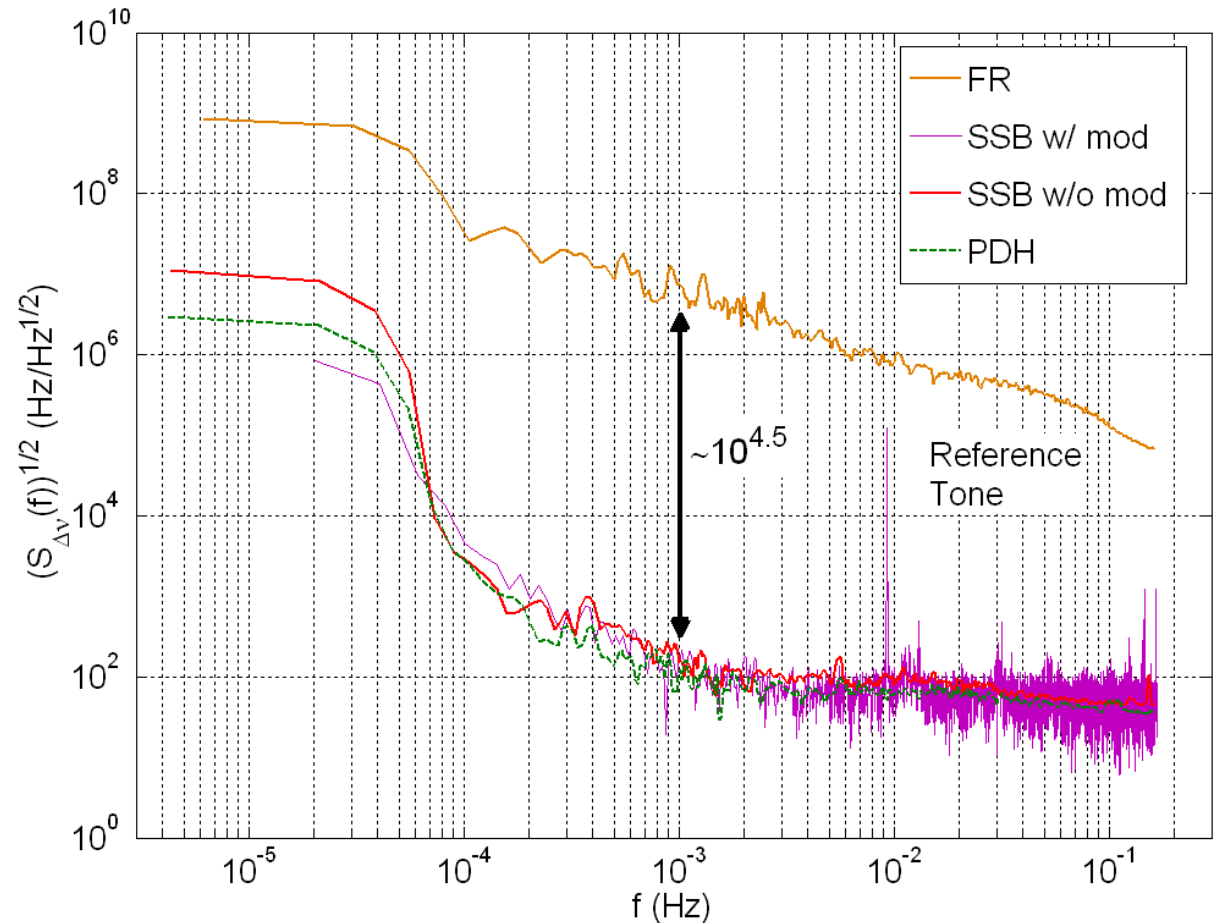


Sideband Lock



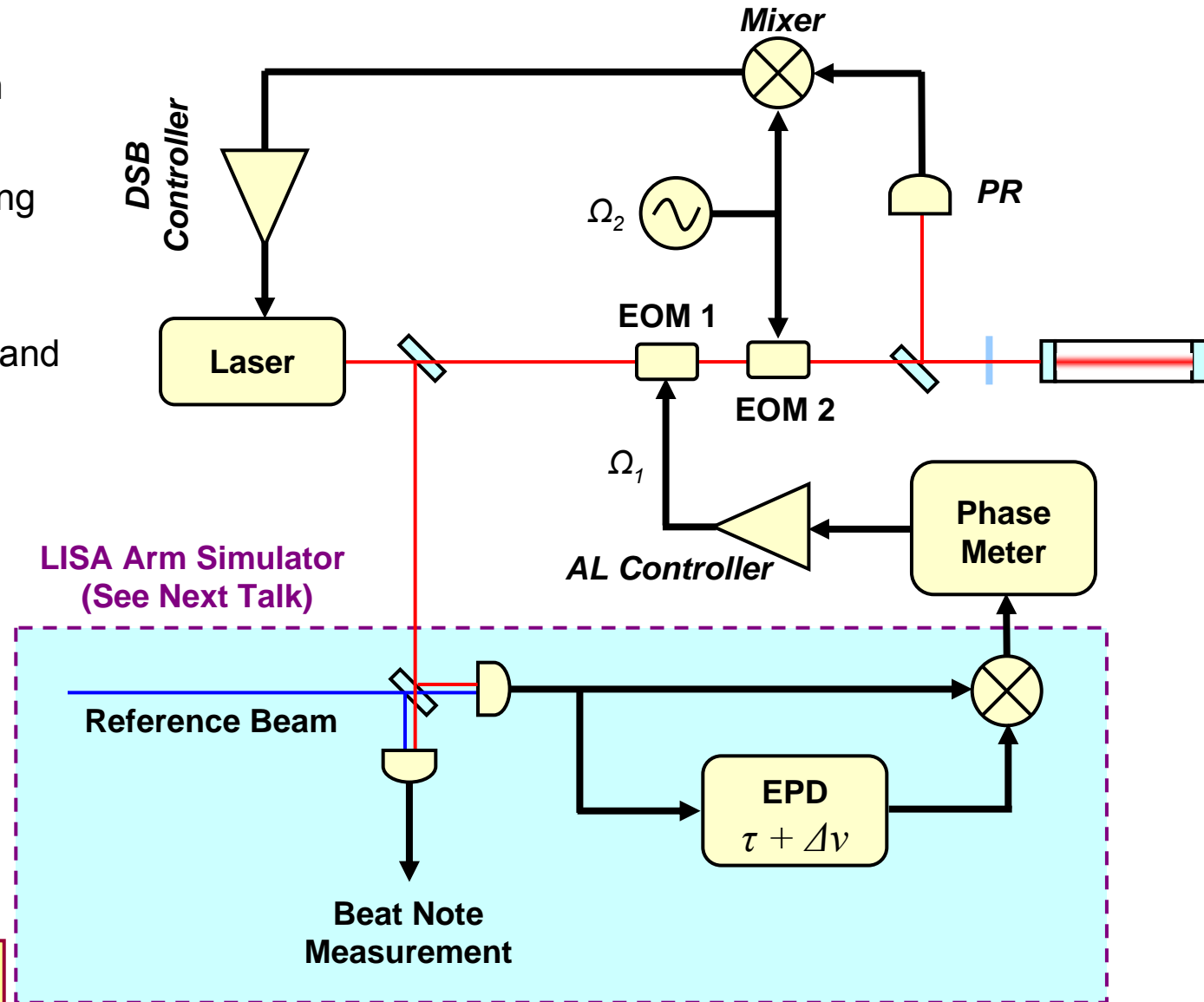
Thorpe, Numata, Livas

- Standard PDH and sideband locking have identical noise performance
- Common technical noises limit both systems.
- Adding modulation tone does not disturb the broadband noise floor.



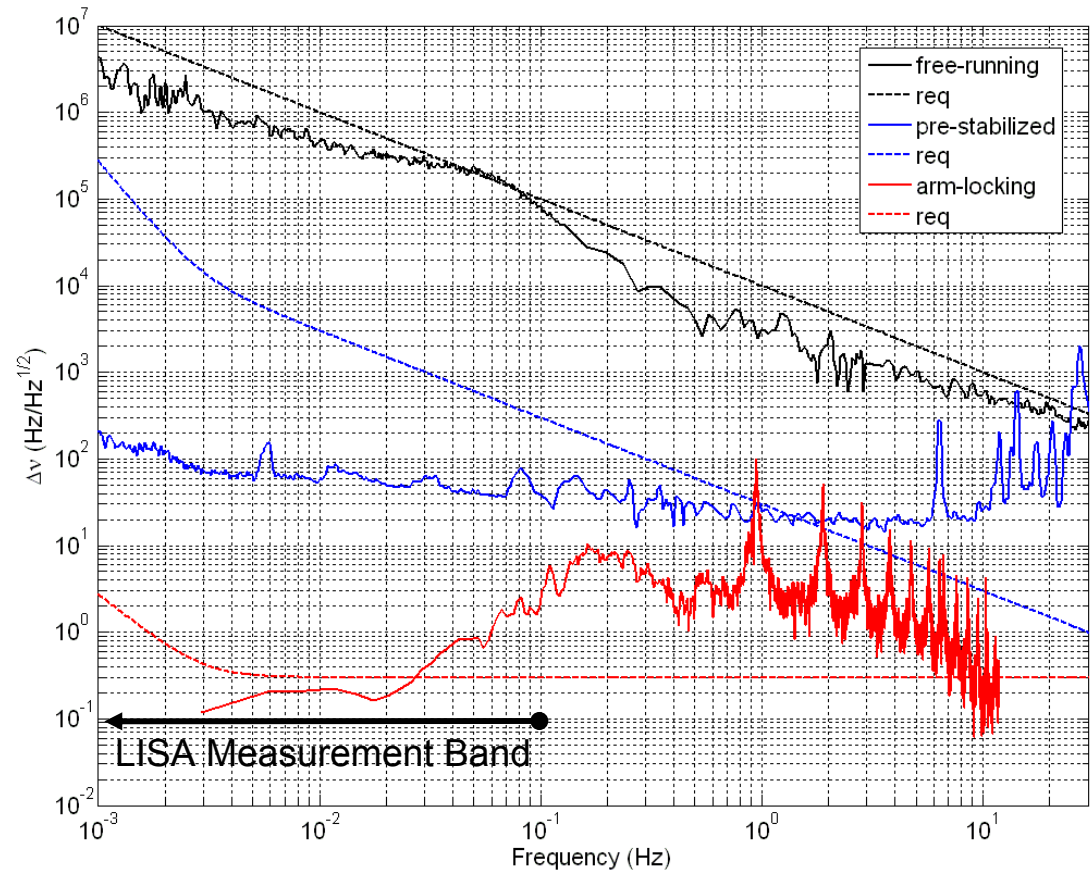
Thorpe, Numata, Livas

- Simulate 1-s long arm using EPD technique
- Pre-stabilize laser using offset sideband locking technique
- Arm-Lock using sideband offset as frequency actuator



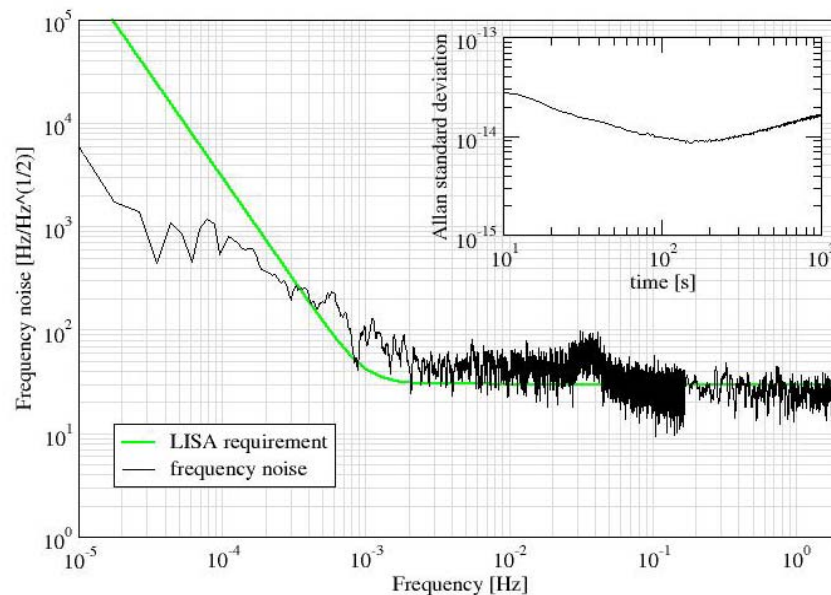
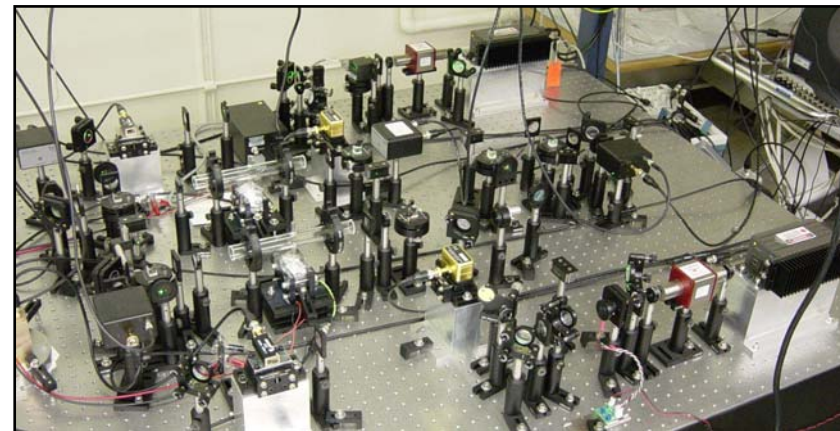
Thorpe, Mitryk, Wand

- Free-running and pre-stabilized lasers *meet LISA requirements in band*.
- Arm-locking system behaves as predicted. (noise spikes at n/τ frequencies)
- Progress towards demonstration of 2/3 of LISA frequency mitigation plan.



Thorpe, Mitryk, Wand

- Spectroscopic reference provides **Absolute reference frequency**
- Laboratory study of frequency stability using two independent Nd:YAG lasers stabilized to hyperfine transition in I_2
- Slightly worse than cavities for $f > 1\text{mHz}$
- Better performance below 0.1 mHz



Leonhart & Camp

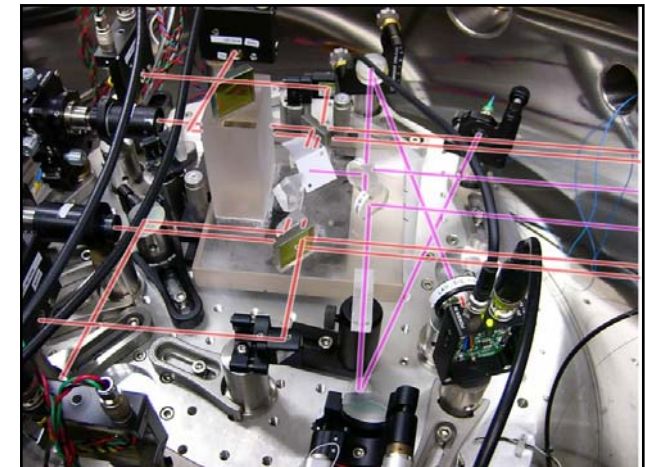
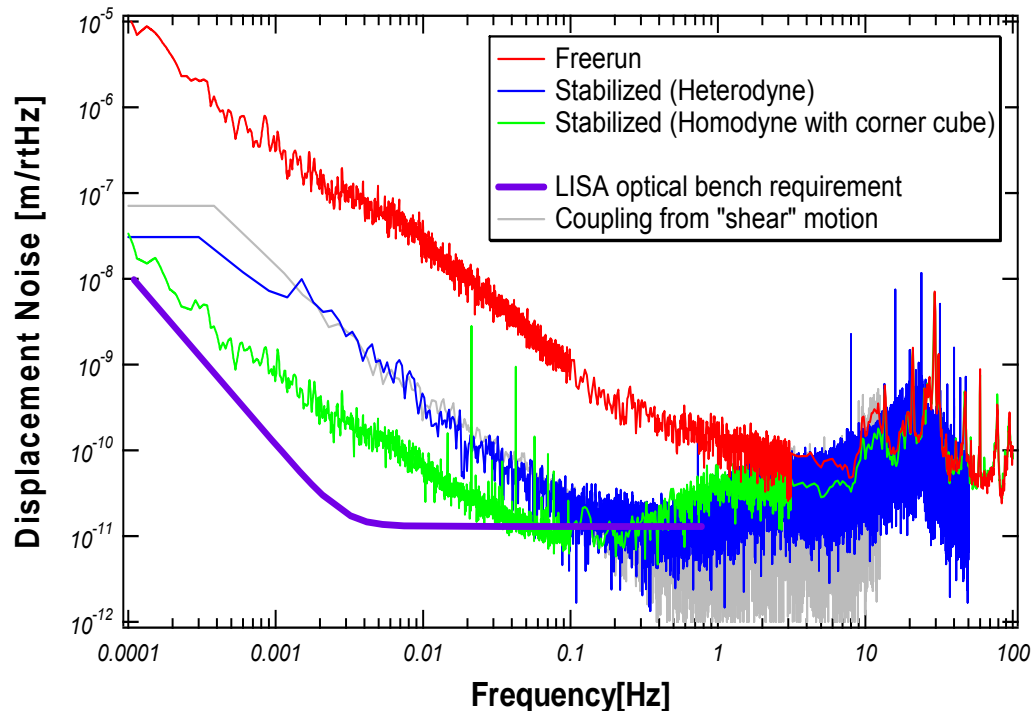
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Testing LISA's inter-spacecraft interferometer on stable platforms

- 2 optical benches with 2 independent pre-stabilized lasers
 - Silicate bonded optical bench, heterodyne interferometer with phasemeter
- 2 degree-of-freedom active control
 - Intended to kill unwanted ground & thermal motion.
 - PZT-based hexapod provides actuation capability.
 - Noise suppression factor: 100~500
 - Performance limited by mechanical coupling from uncontrolled other 4 DoFs.

Numata & Camp



For lowering fundamental noise limit of torsion pendulum

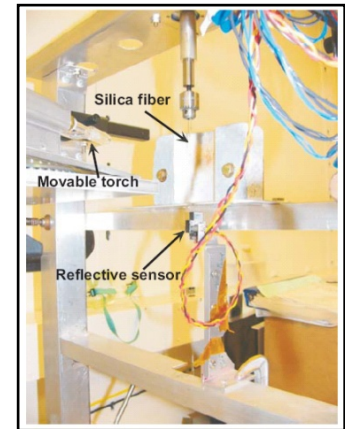
– Our methodology

- Fiber puller, coater, pendulum for loss measurement
- Thin coating technique development

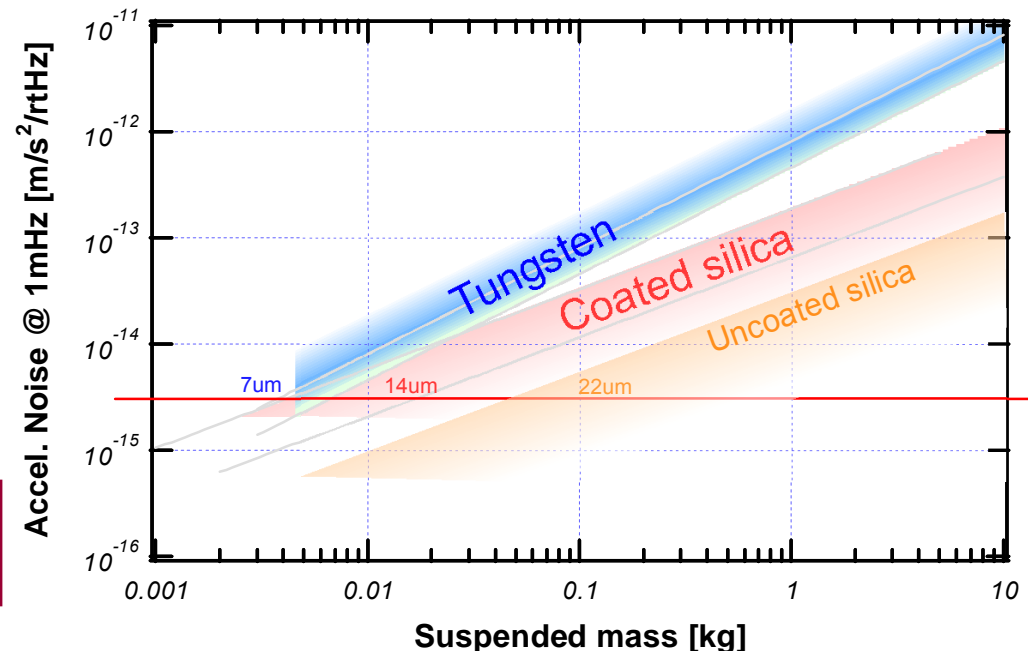
Significant advantages confirmed

– LISA requirement should be reachable with silica

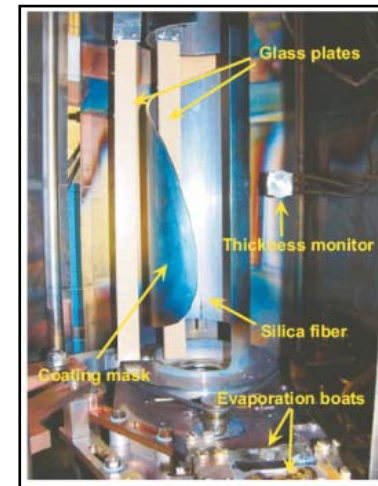
- Test started in LISA torsion pendula in Univ. of Trento & Univ. of Washington



Fiber puller



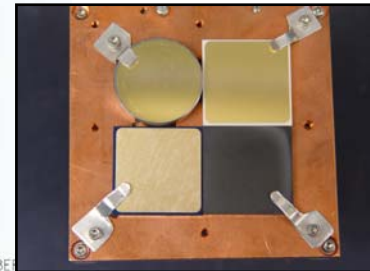
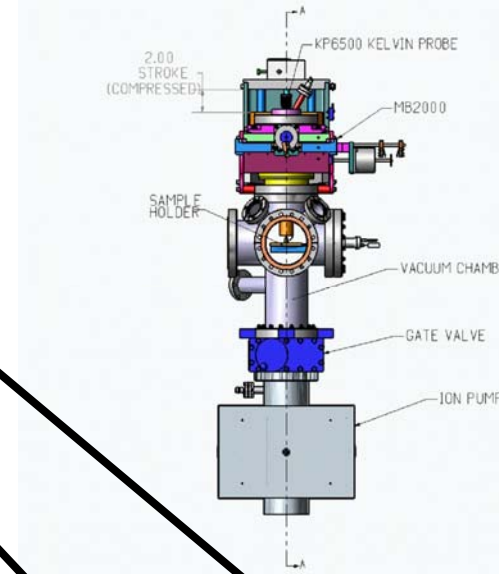
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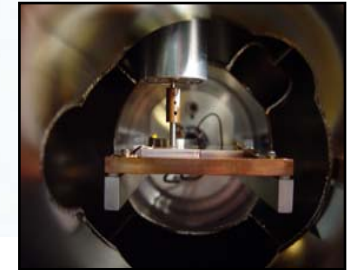
Fiber coater

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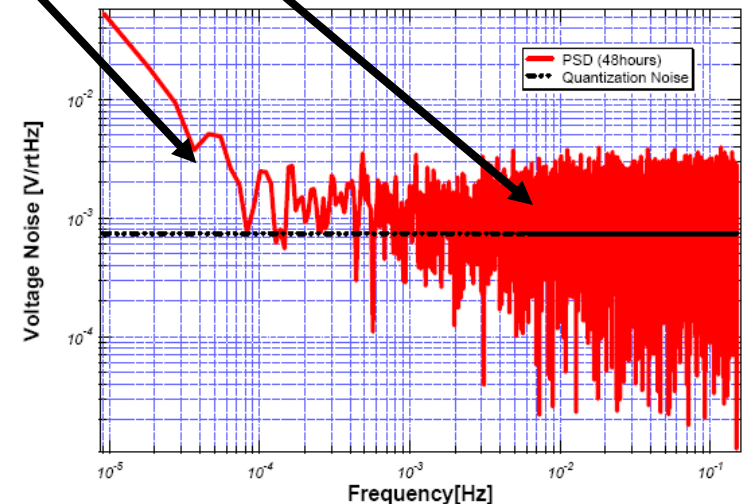
- Vibrating probe induces current proportional to surface potential
- KP limited by ADC quantization noise (recently upgraded)
- Excess low frequency voltage noise of gold surface measured with KP
- Magnitude barely OK for LISA, but cause unknown
- LISA Advantages for patch-effect problem
 - Gold coatings are non-reactive
 - Test mass kept at room temperature







Coated Samples

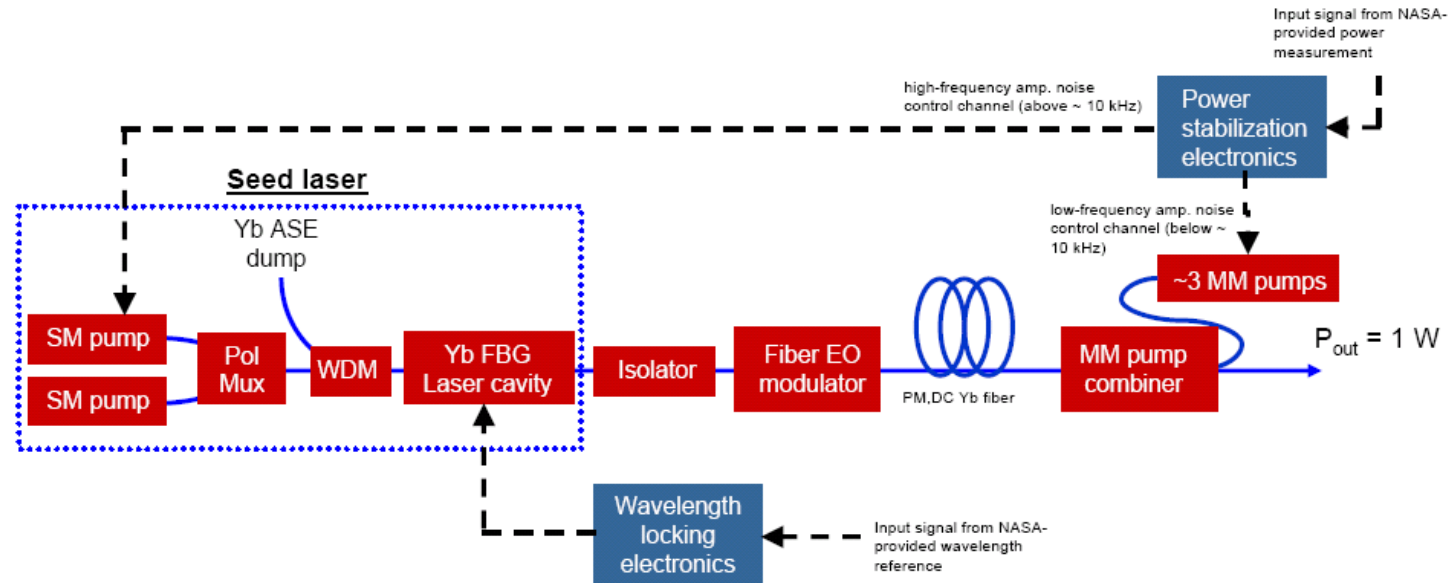


ST-7 sample under test



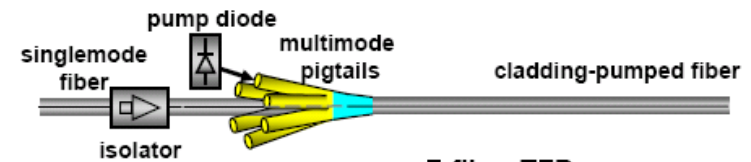
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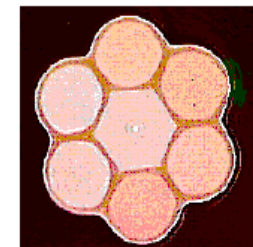


Testing of pump combiner

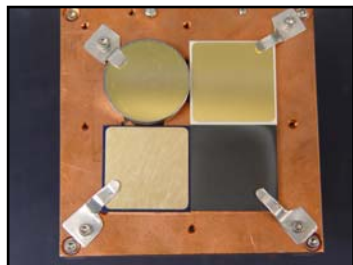
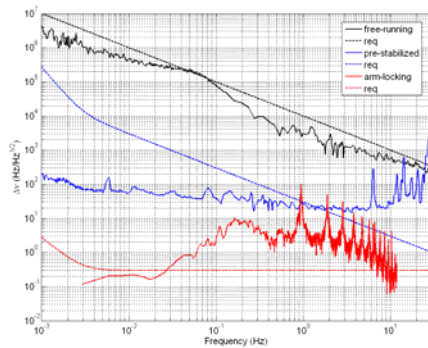
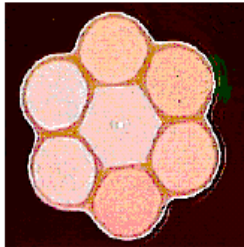
- optical characterization (insertion loss and PER stability) from 5 - 70 C
- thermal screening under high power in vacuum
- temperature cycling in air



7-fiber TFB



J. Camp



Contributors

Jordan Camp

Volker Leonhart

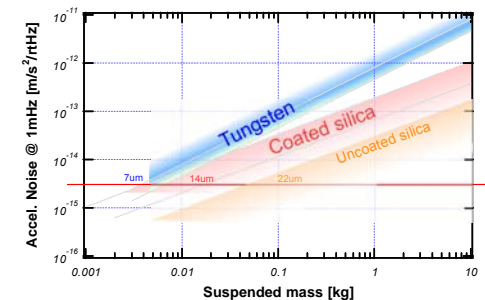
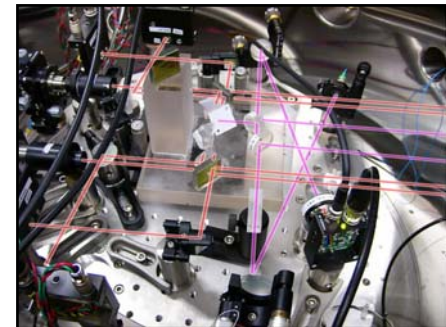
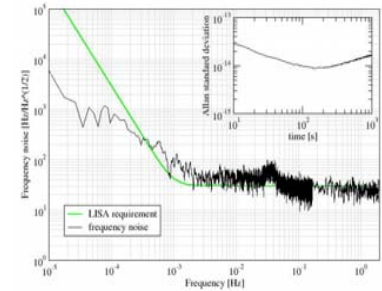
Jeff Livas

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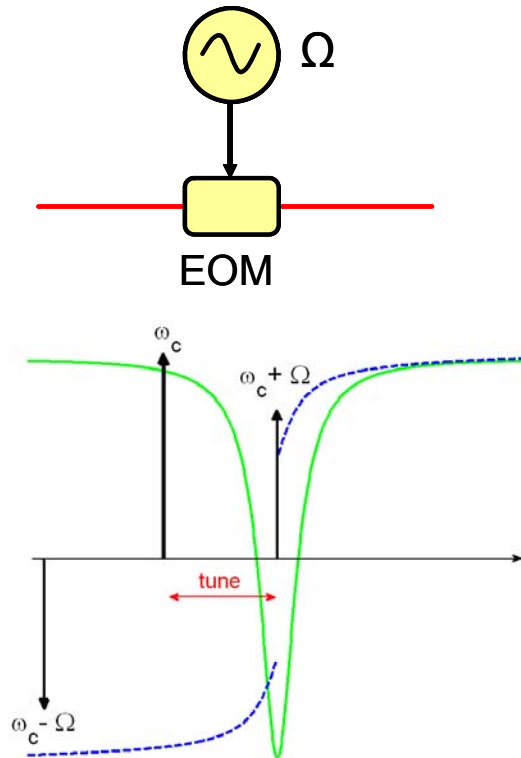




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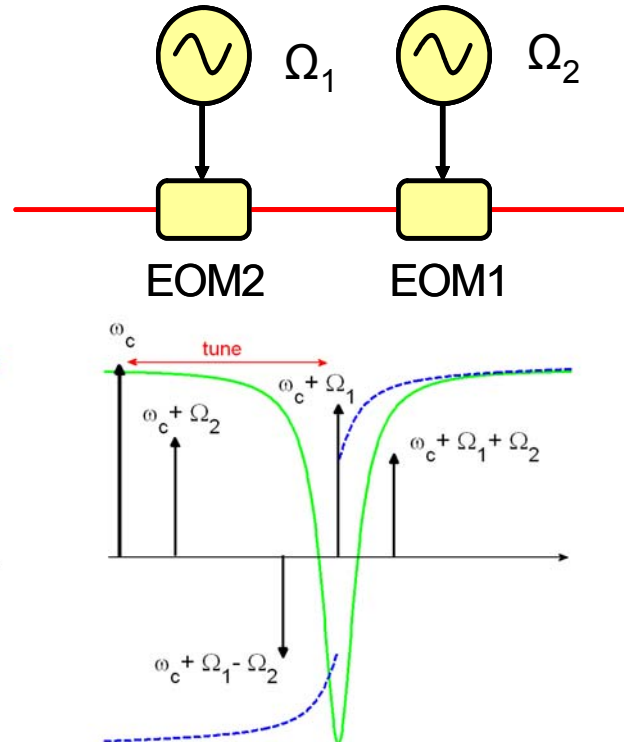
Backup Slides

Single Sideband (SSB)



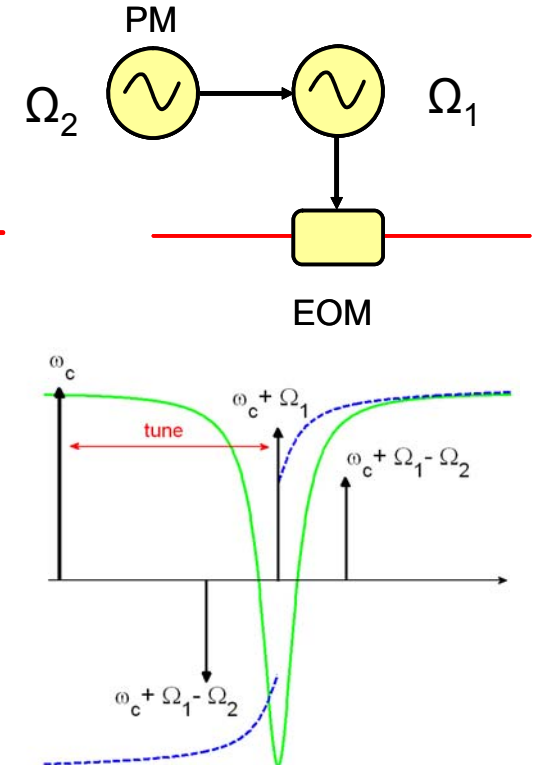
- Simplest to implement
- Some noise coupling due to asymmetry

Dual Sideband (DSB)



- Restores PDH symmetry
- Complex modulation pattern

Electronic Sideband (ESB)



- Simple, symmetric modulation pattern
- Requires phase modulation capability on LO



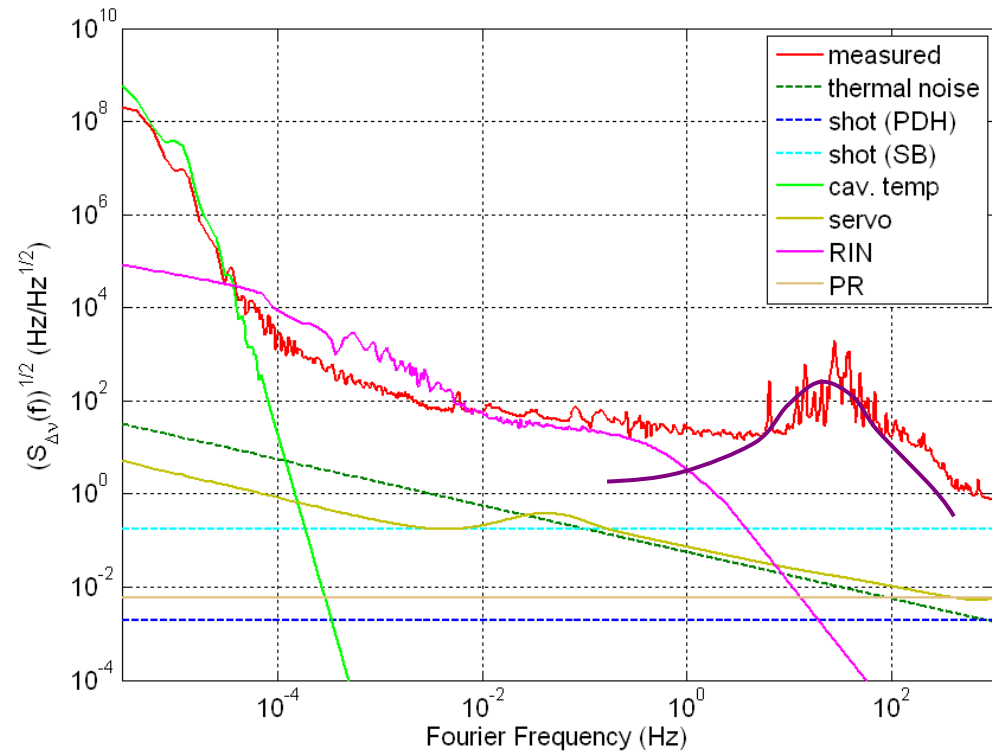
Fundamental Noise

- Shot noise
- Cavity thermal noise



Technical Noise

- Temperature Fluctuations
- Servo Noise
- Photoreceiver noise
- RIN
 - via RFAM
 - via absorption
- Vibration Noise/Acoustic
- Pointing
- ???





Arm-Locking Transfer Function

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- Measured noise suppression matches expectations
- ~40dB at 100mHz

